



Particle motion in Hamiltonian Formalism II Or how to derive and solve equations of

Or how to derive and solve equations of motion

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Summary of Lecture I



- 2nd order dif. equations of motion from Newton's law (in configuration space) can be solved by transforming them to pairs of 1st order dif. equations (in phase space)
- Natural appearance of invariant of motion ("energy")
- Non-linear oscillators have **frequencies** which **depend** on the **invariant** (or "**amplitude**")
- Connected invariant of motion to system's Hamiltonian (derived through Lagrangian)
- Shown that through the Hamiltonian, the equations of motions can be derived
- Poisson bracket operators are helpful for discovering integrals of motion





Canonical transformations



Canonical Transformations



- □ Find a **function** for transforming the Hamiltonian from variable (\mathbf{q}, \mathbf{p}) to (\mathbf{Q}, \mathbf{P}) , so system becomes **simpler** to study
- ☐ Transformation should be **canonical** (or **symplectic**), so that **Hamiltonian** properties (**phase-space volume**) are preserved



Canonical Transformations (CERN)



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- ☐ Transformation should be **canonical** (or **symplectic**), so that Hamiltonian properties (phase-space volume) are preserved
- ☐ These "mixed variable" **generating** functions are derived by

$$F_{1}(\mathbf{q}, \mathbf{Q}) : p_{i} = \frac{\partial F_{1}}{\partial q_{i}}, P_{i} = -\frac{\partial F_{1}}{\partial Q_{i}} F_{3}(\mathbf{Q}, \mathbf{p}) : q_{i} = -\frac{\partial F_{3}}{\partial p_{i}}, P_{i} = -\frac{\partial F_{3}}{\partial Q_{i}}$$

$$F_{2}(\mathbf{q}, \mathbf{P}) : p_{i} = \frac{\partial F_{2}}{\partial q_{i}}, Q_{i} = \frac{\partial F_{2}}{\partial P_{i}} F_{4}(\mathbf{p}, \mathbf{P}) : q_{i} = -\frac{\partial F_{4}}{\partial p_{i}}, Q_{i} = \frac{\partial F_{4}}{\partial P_{i}}$$

A general **non-autonomous** Hamiltonian is transformed to

$$H(\mathbf{Q}, \mathbf{P}, t) = H(\mathbf{q}, \mathbf{p}, t) + \frac{\partial F_j}{\partial t}, \quad j = 1, 2, 3, 4$$



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 \square A general **non-autonomous Hamiltonian** is transformed to

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☐ One generating function can be constructed by the other through **Legendre transformations**, e.g.

$$F_2(\mathbf{q}, \mathbf{P}) = F_1(\mathbf{q}, \mathbf{Q}) - \mathbf{Q} \cdot \mathbf{P}$$
, $F_3(\mathbf{Q}, \mathbf{p}) = F_1(\mathbf{q}, \mathbf{Q}) - \mathbf{q} \cdot \mathbf{p}$, ... with the inner product define as $\mathbf{q} \cdot \mathbf{p} = \sum q_i p_i$



Preservation of Phase Volume



- ☐ A fundamental property of canonical transformations is the **preservation** of **phase space volume**
- ☐ This **volume** preservation in phase space can be represented in the **old** and **new variables** as

$$\int \prod_{i=1}^{n} dp_i dq_i = \int \prod_{i=1}^{n} dP_i dQ_i$$



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☐ The volume element in old and new variables are related through the **Jacobian**

$$\prod_{i=1}^{n} dp_i dq_i = \frac{\partial (P_1, \dots, P_n, Q_1, \dots, Q_n)}{\partial (p_1, \dots, p_n, q_1, \dots, q_n)} \prod_{i=1}^{n} dP_i dQ_i$$



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☐ These two relationships imply that the **Jacobian** of a canonical transformation should have determinant equal to

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$$\left| \frac{\partial (P_1, \dots, P_n, Q_1, \dots, Q_n)}{\partial (p_1, \dots, p_n, q_1, \dots, q_n)} \right| = \left| \frac{\partial (p_1, \dots, p_n, q_1, \dots, q_n)}{\partial (P_1, \dots, P_n, Q_1, \dots, q_n)} \right| = 1$$



Examples of transformations



□ The transformation Q = -p, P = q, which **interchanges conjugate variables** is area preserving, as the Jacobian is

$$\frac{\partial(P,Q)}{\partial(p,q)} = \begin{vmatrix} \frac{\partial P}{\partial p} & \frac{\partial Q}{\partial p} \\ \frac{\partial P}{\partial q} & \frac{\partial Q}{\partial q} \end{vmatrix} = \begin{vmatrix} 0 & -1 \\ 1 & 0 \end{vmatrix} = 1$$





Examples of transformations



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lacktriangle On the other hand, the transformation from **Cartesian to polar** coordinates $q = P \cos Q$, $p = P \sin Q$ is not, since

$$\frac{\partial(q,p)}{\partial(Q,P)} = \begin{vmatrix} -P\sin Q & P\cos Q\\ \cos Q & \sin Q \end{vmatrix} = -P$$





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☐ There are actually "polar" coordinates that are canonical, given by $q = -\sqrt{2P}\cos Q$, $p = \sqrt{2P}\sin Q$ for which

$$\frac{\partial(q,p)}{\partial(Q,P)} = \begin{vmatrix} \sqrt{2P}\sin Q & \sqrt{2P}\cos Q \\ -\frac{\cos Q}{\sqrt{2P}} & \frac{\sin Q}{\sqrt{2P}} \end{vmatrix} = 1$$





The Relativistic Hamiltonian for electromagnetic fields





■ Neglecting self fields and radiation, motion can be described by a "single-particle" Hamiltonian

$$H(\mathbf{x}, \mathbf{p}, t) = c\sqrt{\left(\mathbf{p} - \frac{e}{c}\mathbf{A}(\mathbf{x}, t)\right)^2 + m^2c^2 + e\Phi(\mathbf{x}, t)}$$

$$\mathbf{x} = (x, y, z)$$

$$\mathbf{p} = (p_x, p_y, p_z)$$

$$\Box$$
 $\mathbf{A} = (A_x, A_y, A_z)$

□Ф

Cartesian positions

conjugate momenta

magnetic vector potential

electric scalar potential





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- $\mathbf{x} = (x, y, z)$
- $\mathbf{p} = (p_x, p_y, p_z)$ conjugate momenta
- \Box $\mathbf{A} = (A_x, A_y, A_z)$

Cartesian positions

magnetic vector potential

electric scalar potential

☐ The ordinary kinetic momentum vector is written

$$\mathbf{P} = \gamma m \mathbf{v} = \mathbf{p} - \frac{e}{c} \mathbf{A}$$

with ${\bf v}$ the velocity vector and $\gamma = (1 - v^2/c^2)^{-1/2}$ the relativistic factor





$$H(\mathbf{x}, \mathbf{p}, t) = c\sqrt{\left(\mathbf{p} - \frac{e}{c}\mathbf{A}(\mathbf{x}, t)\right)^2 + m^2c^2 + e\Phi(\mathbf{x}, t)}$$

- ☐ It is generally a 3 degrees of freedom one plus time (i.e., 4 degrees of freedom)
- The Hamiltonian represents the total energy

$$H \equiv E = \gamma mc^2 + e\Phi$$





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- ☐ It is generally a 3 degrees of freedom one plus time (i.e., 4 degrees of freedom)
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$$H \equiv E = \gamma mc^2 + e\Phi$$

☐ The **total kinetic momentum** is

$$P = \left(\frac{H^2}{c^2} - m^2 c^2\right)^{1/2}$$

Using Hamilton's equations

$$(\dot{\mathbf{x}}, \dot{\mathbf{p}}) = [(\mathbf{x}, \mathbf{p}), H]$$

it can be shown that motion is governed by Lorentz equations



From Cartesian to "curved" coordinates



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Particle trajectory

r

☐ It is useful (especially for **rings**) to transform the Cartesian coordinate system to the Frenet-Serret system moving

to a closed curve, with path length S

☐ The **position coordinates** in the two systems are connected by $\mathbf{r} = \mathbf{r_0}(s) + X\mathbf{n}(s) + Y\mathbf{b}(s) = x\mathbf{u_x} + y\mathbf{u_v} + z\mathbf{u_z}$



From Cartesian to "curved" coordinates



Particle trajectory

- ☐ It is useful (especially for **rings**) to transform the Cartesian coordinate system to the **Frenet-Serret system** moving
 - to a closed curve, with path length \dot{S}
- ☐ The **position coordinates** in the two systems are connected by $\mathbf{r} = \mathbf{r_0}(s) + X\mathbf{n}(s) + Y\mathbf{b}(s) = x\mathbf{u_x} + y\mathbf{u_y} + z\mathbf{u_z}$
- □ The **Frenet-Serret unit vectors** and their derivatives are defined as $(\mathbf{t}, \mathbf{n}, \mathbf{b}) = (\frac{d}{ds} \mathbf{r_0}(s), -\rho(s) \frac{d^2}{ds^2} \mathbf{r_0}(s), \mathbf{t} \times \mathbf{n})$

$$\frac{d}{ds} \begin{pmatrix} \mathbf{t} \\ \mathbf{n} \\ \mathbf{b} \end{pmatrix} = \begin{pmatrix} 0 & -\frac{1}{\rho(s)} & 0 \\ \frac{1}{\rho(s)} & 0 & -\tau(s) \\ 0 & 0 & \tau(s) \end{pmatrix} \begin{pmatrix} \mathbf{t} \\ \mathbf{n} \\ \mathbf{b} \end{pmatrix}$$

with $\rho(s)$ the **radius of curvature** and $\tau(s)$ the **torsion** which vanishes in case of planar motion



From Cartesian to "curved" variables



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☐ We are seeking a canonical transformation between

$$(\mathbf{q}, \mathbf{p}) \mapsto (\mathbf{Q}, \mathbf{P}) \text{ or}$$

 $(x, y, z, p_x, p_y, p_z) \mapsto (X, Y, s, P_x, P_y, P_s)$

☐ The **generating** function is

$$(\mathbf{q}, \mathbf{P}) = -(\frac{\partial F_3(\mathbf{p}, \mathbf{Q})}{\partial \mathbf{p}}, \frac{\partial F_3(\mathbf{p}, \mathbf{Q})}{\partial \mathbf{Q}})$$

☐ By using the **relationship** for the **positions**,

$$\mathbf{r} = \mathbf{r_0}(s) + X\mathbf{n}(s) + Y\mathbf{b}(s) = x\mathbf{u_x} + y\mathbf{u_y} + z\mathbf{u_z}$$

the generating function is

$$F_3(\mathbf{p}, \mathbf{Q}) = -\mathbf{p} \cdot \mathbf{r}$$



From Cartesian to "curved" variables



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☐ For planar motion, the momenta are

$$\mathbf{P} = (P_X, P_Y, P_s) = \mathbf{p} \cdot (\frac{\partial F_3}{\partial X}, \frac{\partial F_3}{\partial Y}, \frac{\partial F_3}{\partial s}) = \mathbf{p} \cdot (\mathbf{n}, \mathbf{b}, (1 + \frac{X}{\rho})\mathbf{t})$$

☐ Taking into account that the **vector potential** is also transformed in the same way

$$(A_X, A_Y, A_s) = \mathbf{A} \cdot (\mathbf{n}, \mathbf{b}, (1 + \frac{X}{\rho})\mathbf{t})$$

the **new Hamiltonian** is given by

$$\mathcal{H}(\mathbf{Q}, \mathbf{P}, t) = c \sqrt{(P_X - \frac{e}{c}A_X)^2 + (P_Y - \frac{e}{c}A_Y)^2 + \frac{(P_s - \frac{e}{c}A_s)^2}{(1 + \frac{X}{\rho(s)})^2} + m^2c^2} + e\Phi$$



Changing of the independent variable



- ☐ It is more convenient to use the **path length** *s* , instead of the **time** as **independent variable**
- ☐ The Hamiltonian can be considered as having 4 degrees of freedom, where the 4th "position" is time and its conjugate momentum is $P_t = -\mathcal{H}$



Changing of the independent variable



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- □ The Hamiltonian can be considered as having 4 degrees of freedom, where the 4th "position" is time and its conjugate momentum is $P_t = -\mathcal{H}$
- □ In the same way, the new Hamiltonian with the path length as the independent variable is just $P_s = -\tilde{\mathcal{H}}(X, Y, t, P_X, P_Y, P_t, s)$ with

$$\tilde{\mathcal{H}} = -\frac{e}{c}A_s - \left(1 + \frac{X}{\rho(s)}\right)\sqrt{(\frac{P_t + e\Phi}{c})^2 - m^2c^2 - (P_x - \frac{e}{c}A_X)^2 - (P_Y - \frac{e}{c}A_Y)^2}$$

- ☐ It can be proved that this is indeed a canonical transformation
- Note the existence of the **reference orbit** for **zero vector potential**, for which $(X, Y, P_X, P_Y, P_s) = (0, 0, 0, 0, 0, P_0)_{24}$



Neglecting electric fields



☐ Due to the fact that **longitudinal** (synchrotron) motion is **much slower** than the **transverse** (betatron) one, the electric field can be set to **zero** and the Hamiltonian is written as

$$\tilde{\mathcal{H}} = -\frac{e}{c}A_s - \left(1 + \frac{X}{\rho(s)}\right)\sqrt{(\frac{\mathcal{H}}{c})^2 - m^2c^2} - (P_x - \frac{e}{c}A_X)^2 - (P_Y - \frac{e}{c}A_Y)^2}$$

$$P^2$$



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☐ The Hamiltonian is then written as

$$\tilde{\mathcal{H}} = -\frac{e}{c}A_s - \left(1 + \frac{X}{\rho(s)}\right)\sqrt{(P^2 - (P_x - \frac{e}{c}A_X)^2 - (P_Y - \frac{e}{c}A_Y)^2}$$

☐ If **static** magnetic fields are considered, the time dependence is also dropped, and the system is having **2 degrees of freedom** + "time" (path length)



Momentum rescaling



☐ Due to the fact that **total momentum** is **much larger** than the transverse ones, another transformation may be considered, where the transverse momenta are rescaled

$$(\mathbf{Q}, \mathbf{P}) \mapsto (\bar{\mathbf{q}}, \bar{\mathbf{p}}) \text{ or}$$

$$(X, Y, t, P_X, P_Y, P_t) \mapsto (\bar{x}, \bar{y}, \bar{t}, \bar{p}_x, \bar{p}_y, \bar{p}_t) = (X, Y, -c \ t, \frac{P_X}{P_0}, \frac{P_Y}{P_0}, -\frac{P_t}{P_0c})$$



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☐ The new variables are indeed canonical if the Hamiltonian is also rescaled and written as

$$\begin{split} \bar{\mathcal{H}}(\bar{x},\bar{y},\bar{t},\bar{p}_{x},\bar{p}_{y},\bar{p}_{t}) &= \frac{\tilde{\mathcal{H}}}{P_{0}} = -e\bar{A}_{s} - \left(1 + \frac{\bar{x}}{\rho(s)}\right)\sqrt{\bar{p}_{t}^{2} - \frac{m^{2}c^{2}}{P_{0}}} - (\bar{p}_{x} - e\bar{A}_{x})^{2} - (\bar{p}_{y} - e\bar{A}_{y})^{2} \\ \text{with} \quad \left(\bar{A}_{x},\bar{A}_{y},\bar{A}_{z}\right) &= \frac{1}{P_{0}\,c}(A_{x},A_{y},A_{s}) \end{split}$$

and $\frac{m^2c^2}{P_0} = \frac{1}{\beta_0^2\gamma_0^2}$



Moving the reference frame



 \Box Along the reference trajectory $\bar{p}_{t0} = \frac{1}{\beta_0}$ $\frac{d\bar{t}}{ds}\big|_{P=P_0} = \frac{\partial \bar{H}}{\partial \bar{p}_t}\big|_{P=P_0} = -\bar{p}_{t0} = -\frac{1}{\beta_0}$

☐ It is thus useful to **move** the **reference frame** to the reference trajectory for which another canonical transformation is performed

 $(\bar{\mathbf{q}}, \bar{\mathbf{p}}) \mapsto (\hat{\mathbf{q}}, \hat{\mathbf{p}}) \text{ or }$

$$(\bar{x}, \bar{y}, \bar{t}, \bar{p}_x, \bar{p}_y, \bar{p}_t) \mapsto (\hat{x}, \hat{y}, \hat{t}, \hat{p}_x, \hat{p}_y, \hat{p}_t) = (\hat{x}, \hat{y}, \bar{t} + \frac{s - s_0}{\beta_0}, \hat{p}_x, \hat{p}_y, \bar{p}_t - \frac{1}{\beta_0})$$





Moving the reference frame



 \square Along the reference trajectory $\bar{p}_{t0} = \frac{1}{\beta_0}$ and $\frac{d\bar{t}}{ds}\Big|_{P=P_0} = \frac{\partial \bar{H}}{\partial \bar{p}_t}\Big|_{P=P_0} = -\bar{p}_{t0} = -\frac{1}{\beta_0}$

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$$\Box \text{The mixed variable generating function is}$$

$$(\hat{\mathbf{q}}, \bar{\mathbf{p}}) = (\frac{\partial F_2(\bar{\mathbf{q}}, \hat{\mathbf{p}})}{\partial \hat{\mathbf{p}}}, \frac{\partial F_2(\bar{\mathbf{q}}, \hat{\mathbf{p}})}{\partial \bar{\mathbf{q}}}) \text{ providing}$$

$$F_2(\mathbf{\bar{q}}, \mathbf{\hat{p}}) = \bar{x}\hat{p}_x + \bar{y}\hat{p}_y + (\bar{t} + \frac{s - s_0}{\beta_0})(\hat{p}_t + \frac{1}{\beta_0})$$

$$\Box \text{The Hamiltonian is then}$$

$$\hat{\mathcal{H}}(\hat{x}, \hat{y}, \hat{t}, \hat{p}_x, \hat{p}_y, \hat{p}_t) = \frac{1}{\beta_0} \left(\frac{1}{\beta_0} + \hat{p}_t \right) - e\hat{A}_s - \left(1 + \frac{\hat{x}}{\rho(s)} \right) \sqrt{(\hat{p}_t + \frac{1}{\beta_0})^2 - \frac{1}{\beta_0^2 \gamma_0^2} - (\hat{p}_x - e\hat{A}_x)^2 - (\hat{p}_y - e\bar{A}_y)^2}$$





Relativistic and transverse field approximations



- First note that $\hat{p}_t = \bar{p}_t \frac{1}{\beta_0} = \bar{p}_t \bar{p}_{t0} = \frac{P_t P_0}{P_0} \equiv \delta$ and $l = \hat{t}$
- lacksquare In the ultra-relativistic limit $eta_0 o 1 \;,\;\; rac{1}{eta_0^2 \gamma^2} o 0$ and the Hamiltonian is written as

and the Hamiltonian is written as
$$\rho_0^{\gamma\gamma}$$
 and the Hamiltonian is written as
$$\mu(x,y,l,p_x,p_y,\delta) = (1+\delta) - e\hat{A}_s - \left(1 + \frac{x}{\rho(s)}\right)\sqrt{(1+\delta)^2 - (p_x - e\hat{A}_x)^2 - (p_y - e\hat{A}_y)^2}$$
 where the "hats" are dropped for simplicity



Relativistic and transverse field approximations



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- In the **ultra-relativistic limit** $\beta_0 \to 1$, $\frac{1}{\beta_0^2 \gamma^2} \to 0$ and the Hamiltonian is written as

$$\mathcal{H}(x,y,l,p_x,p_y,\delta) = (1+\delta) - e\hat{A}_s - \left(1 + \frac{x}{\rho(s)}\right)\sqrt{(1+\delta)^2 - (p_x - e\hat{A}_x)^2 - (p_y - e\hat{A}_y)^2}$$

where the "hats" are dropped for simplicity

☐ If we consider **only transverse field** components, the **vector potential** has **only** a **longitudinal** component and the Hamiltonian is written as

$$\mathcal{H}(x,y,l,p_x,p_y,\delta) = (1+\delta) - e\hat{A}_s - \left(1 + \frac{x}{\rho(s)}\right)\sqrt{(1+\delta)^2 - p_x^2 - p_y^2}$$

□ Note that the Hamiltonian is non-linear even in the absence of any field component (i.e. for a drift)!



Canonical transformations and approximations



- ■Summary of canonical transformations and approximations
 - □ From Cartesian to Frenet-Serret (rotating) coordinate system (bending in the horizontal plane)
 - Changing the independent variable from time to the path length s
 - Electric field set to zero, as longitudinal (synchrotron) motion is much slower then transverse (betatron) one
 - Consider static and transverse magnetic fields
 - Rescale the momentum and move the origin to the periodic orbit
 - \square For the ultra-relativistic limit $\beta_0 \to 1$, $\frac{1}{\beta_0^2 \gamma^2} \to 0$ the Hamiltonian becomes

$$\mathcal{H}(x,y,l,p_x,p_y,\delta) = (1+\delta) - e\hat{A}_s - \left(1 + \frac{x}{\rho(s)}\right)\sqrt{(1+\delta)^2 - p_x^2 - p_y^2}$$
 with
$$\frac{P_t - P_0}{P_0} \equiv \delta$$



High-energy, large ring approximation



- ☐ It is useful for study purposes (especially for finding an "integrable" version of the Hamiltonian) to make an extra approximation
- ☐ For this, **transverse momenta** (rescaled to the reference momentum) are considered to be **much smaller** than **1**, i.e. the square root can be expanded.
- \square Considering also the large machine approximation $x << \rho$, (dropping cubic terms), the Hamiltonian is simplified to

$$\mathcal{H} = \frac{p_x^2 + p_y^2}{2(1+\delta)} - \frac{x(1+\delta)}{\rho(s)} - e\hat{A}_s$$

☐ This expansion may **not** be **a good idea**, especially for **low energy**, **small** size **rings**



Linear magnetic fields



Assume a simple case of linear transverse magnetic

$$B_x = b_1(s)y$$

$$B_y = -b_0(s) + b_1(s)x$$

- main bending field
- normalizedquadrupole gradient
- magnetic rigidity

$$-B_0 \equiv b_0(s) = \frac{P_0 c}{e\rho(s)} [T]$$

$$K(s) = b_1(s) \frac{e}{cP_0} = \frac{b_1(s)}{B\rho} [1/\text{m}^2]$$

$$B\rho = \frac{P_0c}{e} [T \cdot m]$$



Linear magnetic fields



Assume a simple case of linear transverse magnetic

fields,
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- magnetic rigidity

- $-B_0 \equiv b_0(s) = \frac{P_0 c}{e \rho(s)}$ [T] $K(s) = b_1(s) \frac{e}{cP_0} = \frac{b_1(s)}{B\rho} [1/\text{m}^2]$
- $B\rho = \frac{P_0c}{1} [T \cdot m]$
- The vector potential has only a longitudinal component which in curvilinear coordinates is

$$B_x = -\frac{1}{1+\frac{x}{o(s)}} \frac{\partial A_s}{\partial y} , \quad B_y = \frac{1}{1+\frac{x}{o(s)}} \frac{\partial A_s}{\partial x}$$

The previous expressions can be integrated to give

$$A_s(x,y,s) = \frac{P_0c}{e} \left[-\frac{x}{\rho(s)} - \left(\frac{1}{\rho(s)^2} + K(s) \right) \frac{x^2}{2} + K(s) \frac{y^2}{2} \right] = P_0c \ \hat{A}_s(x,y,s)$$



The integrable Hamiltonian



The Hamiltonian for linear fields can be finally written as

$$\mathcal{H} = \frac{p_x^2 + p_y^2}{2(1+\delta)} - \frac{x\delta}{\rho(s)} + \frac{x^2}{2\rho(s)^2} + \frac{K(s)}{2}(x^2 - y^2)$$

Hamilton's equation are $\frac{\frac{dx}{ds} = \frac{p_x}{1+\delta}}{\frac{dy}{ds}} = \frac{\delta}{\rho(s)} - \left(\frac{1}{\rho^2(s)} + K(s)\right)x$

and they can be written as two second order uncoupled differential equations, i.e. **Hill's equations**

$$x'' + \frac{1}{1+\delta} \left(\frac{1}{\rho(s)^2} + K(s) \right) x = \frac{\delta}{\rho(s)}$$
 with the usual solution for
$$y'' - \frac{1}{1+\delta} K(s) y = 0$$

$$\delta = 0 \text{ and } u = x, y$$

$$u(s) = \sqrt{\epsilon \beta(s)} \cos(\psi(s) + \psi_0)$$

$$u'(s) = \sqrt{\frac{\epsilon}{\beta(s)}} (\sin(\psi(s) + \psi_0) + \alpha(s) \cos(\psi(s) + \psi_0))$$



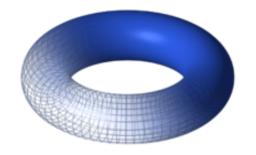
Action-angle variables



- There is a canonical transformation to some **optimal** set of variables which can simplify the phase-space motion
- This set of variables are the **action-angle** variables
- The action vector is defined as the integral $\mathbf{J} = \oint \mathbf{p} d\mathbf{q}$ over closed paths in phase space.
- An integrable Hamiltonian is written as a function of only the actions, i.e. $H_0 = H_0(\mathbf{J})$. Hamilton's equations give

$$\dot{\phi}_i = \frac{\partial H_0(\mathbf{J})}{\partial J_i} = \omega_i(\mathbf{J}) \Rightarrow \phi_i = \omega_i(\mathbf{J})t + \phi_{i0}$$

$$\dot{J}_i = -\frac{\partial H_0(\mathbf{J})}{\partial \phi_i} = 0 \Rightarrow J_i = \text{const.}$$



- i.e. the actions are integrals of motion and the angles are evolving linearly with time, with constant frequencies which depend on the actions
- The actions define the surface of an invariant torus, topologically equivalent to the product of n circles



Accelerator Hamiltonian in action-angle variables



Considering on-momentum motion, the Hamiltonian can be written as

$$\mathcal{H} = \frac{p_x^2 + p_y^2}{2} + \frac{K_x(s)x^2 - K_y(s)y^2}{2}$$

■ The generating function from the original to action angle variables is

$$F_1(x, y, \phi_x, \phi_y; s) = -\frac{x^2}{2\beta_x(s)} \left[\tan \phi_x(s) + a_x(s) \right] - \frac{y^2}{2\beta_y(s)} \left[\tan \phi_y(s) + a_y(s) \right]$$

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■ The old variables with respect to actions and angles are

$$u(s) = \sqrt{2\beta_u(s)J_u}\cos\phi_u(s)\;,\;\; p_u(s) = -\sqrt{\frac{2J_u}{\beta_u(s)}}\left(\sin\phi_u(s) + \alpha_u(s)\cos\phi_u(s)\right)$$
 and the Hamiltonian takes the form

$$\mathcal{H}_0(J_x,J_y,s)=rac{J_x}{eta_x(s)}+rac{J_y}{eta_y(s)}$$

■ The "time" (longitudinal position) dependence can be eliminated by the transformation to **normalized coordinate**

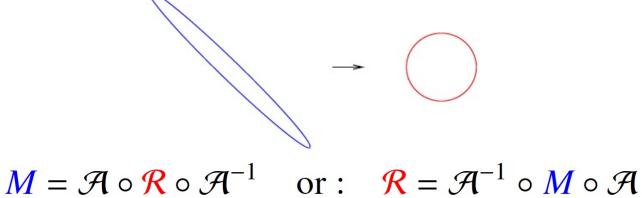
$$\begin{pmatrix} \mathcal{U} \\ \mathcal{U}' \end{pmatrix} = \begin{pmatrix} \frac{1}{\sqrt{\beta}} & 0 \\ \frac{\alpha}{\sqrt{\beta}} & \sqrt{\beta} \end{pmatrix} \begin{pmatrix} u \\ u' \end{pmatrix} \quad \text{or} \begin{pmatrix} \mathcal{U} \\ \mathcal{U}' \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \cos(\nu\phi) \\ \sin(\nu\phi) \end{pmatrix} \text{with } \nu = \frac{1}{2\pi} \oint \frac{du}{\beta(s)}$$



Linear normal forms



Make a coordinate transformation so that we get a simpler form of the matrix, i.e. ellipses are transformed to circles (simple rotation)



Using linear algebra, the solution is

$$\mathcal{A} = \begin{pmatrix} \sqrt{\beta(s_0)} & 0 \\ -\frac{\alpha(s_0)}{\sqrt{\beta(s_0)}} & \frac{1}{\sqrt{\beta(s_0)}} \end{pmatrix} \quad \text{and} \quad \mathcal{R} = \begin{pmatrix} \cos(\mu_x) & \sin(\mu_x) \\ -\sin(\mu_x) & \cos(\mu_x) \end{pmatrix}$$

■ This transformation can be extended to a non-linear system (see Advanced course)



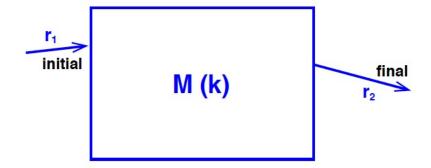


Symplectic maps





A generalization of the matrix (which can only describe linear systems), is a map, which transforms a system from some initial to some final coordinates



- Analyzing the map, will give useful information about the behavior of the system
- There are different ways to build the map:
 - □ Taylor (Power) maps
 - Lie transformations
 - □ Truncated Power Series Algebra (TPSA), can generate maps from straight-forward tracking
- Preservation of symplecticity is important



Symplectic maps



- Consider two sets of canonical variables Z, may be even considered as the evolution of the system between two points in phase space
- A transformation from the one to the other set can be constructed through a map $\mathcal{M}: \mathbf{z} \mapsto \mathbf{\bar{z}}$
- lacksquare The **Jacobian matrix** of the map $M=M(\mathbf{z},t)$ is composed by the elements $M_{ij} \equiv \frac{\partial \bar{z}_i}{\partial z_i}$
- The map is **symplectic** if $M^TJM = J$ where $J = \begin{pmatrix} \mathbf{0} & \mathbf{I} \\ -\mathbf{I} & \mathbf{0} \end{pmatrix}$ It can be shown that $\det(M) = 1$
- It can be shown that the variables defined through a symplectic map $[\bar{z}_i, \bar{z}_j] = [z_i, z_j] = J_{ij}$ which is a known relation satisfied by canonical variables
- In other words, symplectic maps **preserve** Poisson brackets



Why symplecticity is important



- Symplecticity guarantees that the transformations in phase space are area preserving
- To understand what deviation from symplecticity produces consider the simple case of the **quadrupole** with the general matrix written as

$$\mathcal{M}_{Q} = \begin{pmatrix} \cos(\sqrt{k}L) & \frac{1}{\sqrt{k}}\sin(\sqrt{k}L) \\ -\sqrt{k}\sin(\sqrt{k}L) & \cos(\sqrt{k}L) \end{pmatrix}$$

■ Take the Taylor expansion for small lengths, up to first order

$$\mathcal{M}_{\mathbf{Q}} = \begin{pmatrix} 1 & L \\ -kL & 1 \end{pmatrix} + O(L^2)$$

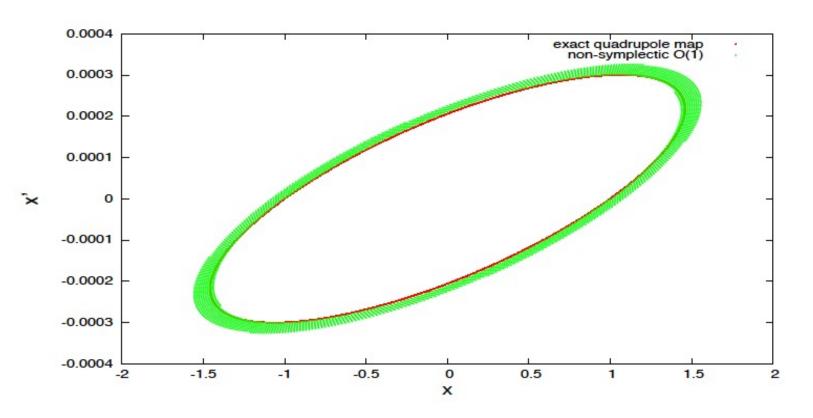
This is indeed **not symplectic** as the determinant of the matrix is equal to $1+kL^2$, i.e. there is a deviation from symplecticity at $2^{\rm nd}$ order in the quadrupole length



Phase portrait for non-symplectic matrix



- The iterated non-symplectic matrix does not provide the well-know elliptic trajectory in phase space
- Although the trajectory is very close to the original one, it spirals outwards towards infinity





Lie formalism



- The Poisson bracket properties satisfy what is mathematically called a **Lie** algebra
- They can be represented by (Lie) operators of the form

$$f: g = [f, g]$$
 and $f: f: {}^2g = [f, [f, g]]$ etc.



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- For a Hamiltonian system $H(\mathbf{z},t)$ there is a **formal** solution of the equations of motion $\frac{d\mathbf{z}}{dt} = [H,\mathbf{z}] =: H: \mathbf{z}$ written as $\mathbf{z}(t) = \sum_{k=0}^{\infty} \frac{t^k : H : k}{k!} \mathbf{z}_0 = e^{t : H:} \mathbf{z}_0$ with a symplectic map $\mathcal{M} = e^{:H:}$



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- The 1-turn accelerator map can be represented by the composition of the maps of each element $\mathcal{M} = e^{:f_2:} e^{:f_3:} e^{:f_4:} \dots$ where f_i (called the generator) is the Hamiltonian for each element, a polynomial of degree m in the variables z_1, \dots, z_n



General non-linear Accelerator Hamiltonian



- Considering the general expression of the the longitudinal component of the vector potential is (see appendix)
 - ☐ In curvilinear coordinates (curved elements)

$$A_s = (1 + \frac{x}{\rho(s)})B_0 \Re e \sum_{n=0}^{\infty} \frac{b_n + ia_n}{n+1} (x + iy)^{n+1}$$

 $\blacksquare \text{ In Cartesian coordinates } A_s = B_0 \Re e \sum_{s=0}^{\infty} \frac{b_n + ia_n}{n+1} (x+iy)^{n+1}$

with the **multipole coefficients** being written as

$$a_n = \frac{1}{B_0 n!} \frac{\partial^n B_x}{\partial x^n} \Big|_{x=y=0}$$
 and $b_n = \frac{1}{B_0 n!} \frac{\partial^n B_y}{\partial x^n} \Big|_{x=y=0}$

The general non-linear Hamiltonian can be written as

$$\mathcal{H}(x, y, p_x, p_y, s) = \mathcal{H}_0(x, y, p_x, p_y, s) + \sum_{k_x, k_y} h_{k_x, k_y}(s) x^{k_x} y^{k_y}$$

with the **periodic functions** $h_{k_x,k_y}(s) = h_{k_x,k_y}(s+C)$





Magnetic element Hamiltonians



Dipole:

$$H = \frac{x\delta}{\rho} + \frac{x^2}{2\rho^2} + \frac{p_x^2 + p_y^2}{2(1+\delta)}$$

Quadrupole:

$$H = \frac{1}{2}k_1(x^2 - y^2) + \frac{p_x^2 + p_y^2}{2(1+\delta)}$$

Sextupole:

Foole:
$$H = \frac{1}{3}k_2(x^3 - 3xy^2) + \frac{p_x^2 + p_y^2}{2(1+\delta)}$$

Octupole:

$$H = \frac{1}{4}k_3(x^4 - 6x^2y^2 + y^4) + \frac{p_x^2 + p_y^2}{2(1+\delta)}$$



Map for quadrupole



Consider the 1D quadrupole Hamiltonian

$$H = \frac{1}{2}(k_1x^2 + p^2)$$

 \blacksquare For a quadrupole of length L, the map is written as

$$e^{\frac{L}{2}:(k_1x^2+p^2):}$$



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Its application to the transverse variables is

$$e^{-\frac{L}{2}:(k_1x^2+p^2):}x = \sum_{n=0}^{\infty} \left(\frac{(-k_1L^2)^n}{(2n)!}x + L\frac{(-k_1L^2)^n}{(2n+1)!}p\right)$$

$$e^{-\frac{L}{2}:(k_1x^2+p^2):}p = \sum_{n=0}^{\infty} \left(\frac{(-k_1L^2)^n}{(2n)!}p - \sqrt{k_1}\frac{(-k_1L^2)^n}{(2n+1)!}p\right)$$



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$$e^{-\frac{L}{2}:(k_1x^2+p^2):}p = \sum_{n=0}^{\infty} \left(\frac{(-k_1L^2)^n}{(2n)!}p - \sqrt{k_1}\frac{(-k_1L^2)^n}{(2n+1)!}p\right)$$

This finally provides the usual quadrupole matrix

$$e^{-\frac{L}{2}:(k_1x^2+p^2):}x = \cos(\sqrt{k_1}L)x + \frac{1}{\sqrt{k_1}}\sin(\sqrt{k_1}L)p$$

$$e^{-\frac{L}{2}:(k_1x^2+p^2):}p = -\sqrt{k_1}\sin(\sqrt{k_1}L)x + \cos(\sqrt{k_1}L)p$$



Appendix





Magnetic multipole expansion



From Gauss law of magnetostatics, a vector potential exist

$$\nabla \cdot \mathbf{B} = 0 \rightarrow \exists \mathbf{A} : \mathbf{B} = \nabla \times \mathbf{A}$$

- Assuming transverse 2D field, vector potential has only one component A_s . The Ampere's law in vacuum (inside the beam pipe) $\nabla \times \mathbf{B} = 0 \quad \rightarrow \quad \exists V: \quad \mathbf{B} = -\nabla V$
- Using the previous equations, the relations between field components and potentials are

$$B_x = -\frac{\partial V}{\partial x} = \frac{\partial A_s}{\partial y} , \quad B_y = -\frac{\partial V}{\partial y} = -\frac{\partial A_s}{\partial x}_y$$

i.e. Riemann conditions of an analytic function

Exists complex potential of z=x+iy with power series expansion convergent in a circle with radius $|z|=r_c$ (distance from iron yoke)

$$\mathcal{A}(x+iy) = A_s(x,y) + iV(x,y) = \sum_{n} \kappa_n z^n = \sum_{n} (\lambda_n + i\mu_n)(x+iy)^n$$

iron



Multipole expansion II



From the complex potential we can derive the fields

$$B_y + iB_x = -\frac{\partial}{\partial x}(A_s(x,y) + iV(x,y)) = -\sum_{n=1}^{\infty} n(\lambda_n + i\mu_n)(x+iy)^{n-1}$$

• Setting $b_n = -n\lambda_n$, $a_n = n\mu_n$

$$B_y + iB_x = \sum_{n=1}^{\infty} (b_n - ia_n)(x + iy)^{n-1}$$

Define normalized coefficients

$$b'_n = \frac{b_n}{10^{-4}B_0}r_0^{n-1}, \ a'_n = \frac{a_n}{10^{-4}B_0}r_0^{n-1}$$

on a reference radius r_0 , 10^{-4} of the main field to get

$$B_y + iB_x = 10^{-4}B_0 \sum_{n=1}^{30} (b'_n - ia'_n) (\frac{x + iy}{r_0})^{n-1}$$

■ Note: n' = n - 1 is the US convention