Beam Cooling

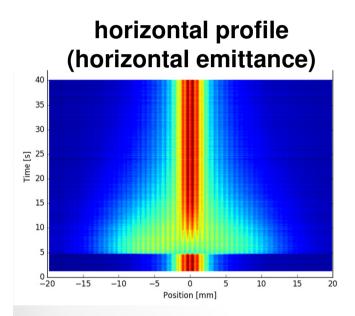
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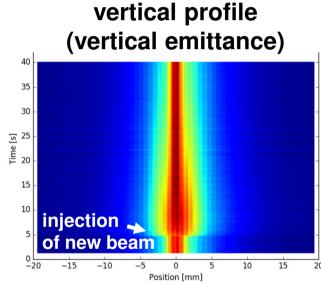
CAS Advanced Accelerator Physics, Royal Holloway University of London, 3 - 15 September 2017

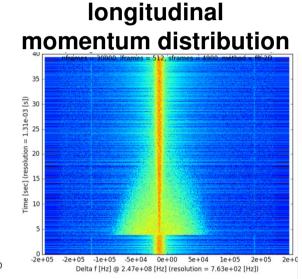
Observation of Cooling

Xe⁵⁴⁺ beam at 400 MeV/u cooled with electron current 200 mA

cooling in six-dimensional phase space







measured with residual gas ionization beam profile monitor

longitudinal Schottky noise

Beam Cooling

Introduction

- 1. Electron Cooling
- 2. Ionization Cooling
- 3. Laser Cooling
- 4. Stochastic Cooling

Beam Cooling

Beam cooling is synonymous for a reduction of beam temperature. Temperature is equivalent to terms as phase space volume, emittance and momentum spread.

Beam Cooling processes are not following Liouville's Theorem: in a system where the particle motion is controlled by external conservative forces the phase space density is conserved' (This neglects interactions between beam particles.)

Beam cooling techniques are non-Liouvillean processes which violate the assumption of a conservative force.

e.g. interaction of the beam particles with other particles (electrons, photons, matter)

Cooling Force

Generic (simplest case of a) cooling force:

$$F_{x,y,s} = -\alpha_{x,y,s} v_{x,y,s}$$

 $v_{x,y,s}$ velocity in the rest frame of the beam

non conservative, cannot be described by a Hamiltonian

For a 2D subspace distribution function f(z, z', t)

$$F_z = -\alpha_z v_z \quad z = x, y, s \quad v_z = v_0 \cdot z'$$

$$\frac{df(z, z', t)}{dt} = -\lambda_z f(z, z', t) \qquad \lambda_z \text{ cooling (damping) rate}$$

in a circular accelerator:

Transverse (emittance) cooling

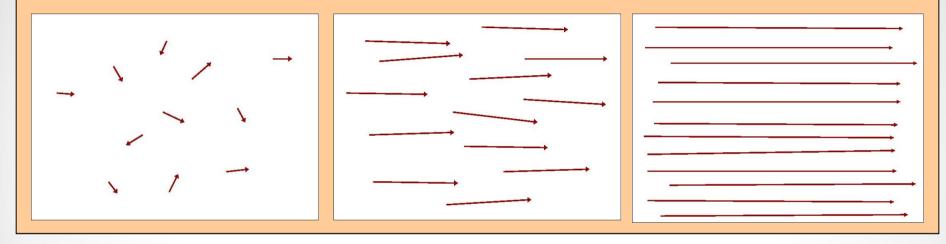
$$\epsilon_{x,y}(t_0+t) = \epsilon_{x,y}(t_0) e^{-\lambda_{x,y}t}$$

Longitudinal (momentum spread) cooling
$$\ \, rac{\delta p_{||}}{p_0}(t_0+t)=rac{\delta p_{||}}{p_0}(t_0)\,\,e^{-\lambda_{||}t}$$

Beam Temperature

Where does the beam temperature originate from?

The beam particles are generated in a 'hot' source



at rest (source)

at low energy

at high energy

In a standard accelerator the beam temperature is not reduced (thermal motion is superimposed the average motion after acceleration)

but: many processes can heat up the beam

e.g. heating by mismatch, space charge, intrabeam scattering, internal targets, residual gas, external noise

Beam Temperature Definition

Longitudinal beam temperature

$$\frac{1}{2}k_B T_{\parallel} = \frac{1}{2}mv_{\parallel}^2 = \frac{1}{2}mc^2\beta^2(\frac{\delta p_{\parallel}}{p})^2$$

Transverse beam temperature

$$\frac{1}{2}k_BT_{\perp} = \frac{1}{2}mv_{\perp}^2 = \frac{1}{2}mc^2\beta^2\gamma^2\theta_{\perp}^2 \qquad \theta_{\perp} = \frac{v_{\perp}}{\beta c}, \quad \theta_{\perp}(s) = \sqrt{\frac{\epsilon}{\beta_{\perp}(s)}}$$

$$\theta_{\perp} = \frac{v_{\perp}}{\beta c}, \quad \theta_{\perp}(s) = \sqrt{\frac{\epsilon}{\beta_{\perp}(s)}}$$

dependent on s

Distribution function

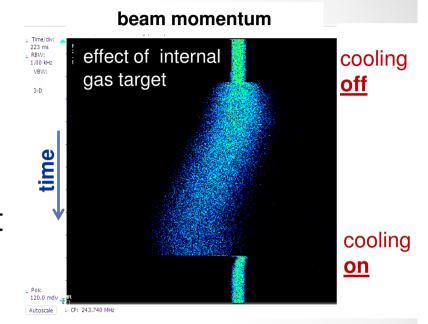
$$f(v_\perp,v_\parallel) \propto \exp(-rac{mv_\perp^2}{2k_BT_\perp} - rac{mv_\parallel^2}{2k_BT_\parallel})$$

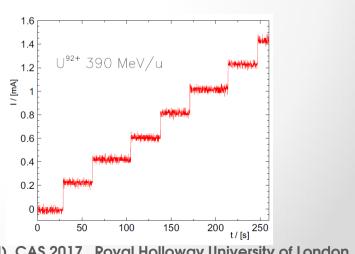
Particle beams can be anisotropic: $k_B T_{||} \neq k_B T_{\perp}$ e.g. due to laser cooling or the distribution of the electron beam

Don't confuse: beam energy ↔ beam temperature (e.g. a beam of energy 100 GeV can have a temperature of 1 eV)

Benefits of Beam Cooling

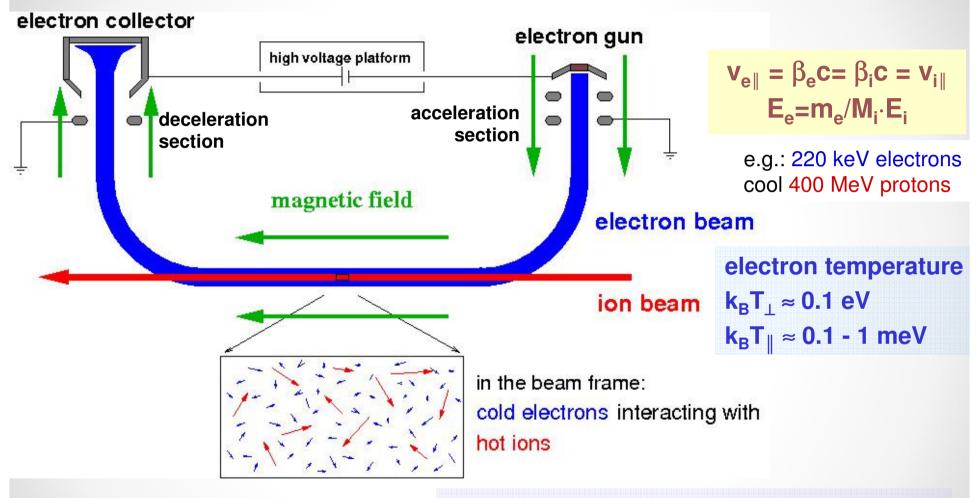
- Improved beam quality
 - Precision experiments
 - Luminosity increase
- Compensation of heating
 - Experiments with internal target
 - Colliding beams
- Intensity increase by accumulation
 - Weak beams from the source can be enhanced
 - Secondary beams (antiprotons, rare isotopes)





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1. Electron Cooling

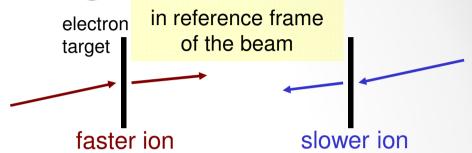


superposition of a cold intense electron beam with the same velocity

momentum transfer by Coulomb collisions cooling force results from energy loss in the co-moving gas of free electrons

Simple Derivation of the Electron Cooling Force

Analogy: energy loss in matter (electrons in the shell)



Rutherford scattering:
$$2 \ tan(\frac{\theta}{2}) = \frac{2Z_1Z_2e^2}{4\pi\epsilon_0\Delta pvb} \quad Z_1 = Q \ (ion), \ Z_2 = -1 \ (electron)$$

 Energy transfer: $\Delta E(b) = \frac{(\Delta p)^2}{2m_e} \simeq \frac{2Q^2e^4}{(4\pi\epsilon_0)^2m_ev^2} \frac{1}{b^2} \ (for \ b \gg b_{min})$

Energy transfer:
$$\Delta E(b)=rac{(\Delta p)^2}{2m_e}\simeqrac{2Q^2e^4}{(4\pi\epsilon_0)^2m_ev^2}rac{1}{b^2}$$
 $(for~b\gg b_{min})$

Minimum impact parameter:
$$b_{min} = \frac{Qe^2}{(4\pi\epsilon_0) \ m_e v^2}$$

from:
$$\Delta E(b_{min}) = \Delta E_{max} \simeq 2m_e v^2$$

Energy loss:

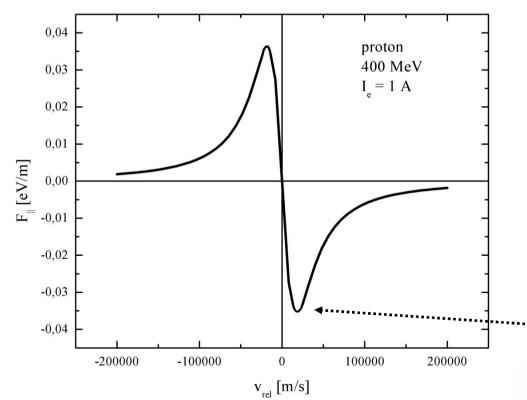
$$-\frac{dE}{dx} = 2\pi \int_{b_{min}}^{b_{max}} b n_e \Delta E \, db = \frac{4\pi Q^2 e^4}{(4\pi\epsilon_0)^2 m_e v^2} n_e \ln \frac{b_{max}}{b_{min}}$$

Coulomb logarithm $L_C=ln (b_{max}/b_{min}) \approx 10$ (typical value)

Characteristics of the Electron Cooling Force

$$\overrightarrow{F}(\overrightarrow{v_i}) = -\frac{4\pi Q^2 e^4 n_e}{(4\pi\epsilon_0)^2 m_e} \int L_C(\overrightarrow{v}_{rel}) f(\overrightarrow{v_e}) \frac{\overrightarrow{v}_{rel}}{v_{rel}^3} d^3 \overrightarrow{v_e}$$

$$\overrightarrow{v}_{rel} = \overrightarrow{v_i} - \overrightarrow{v_e}$$

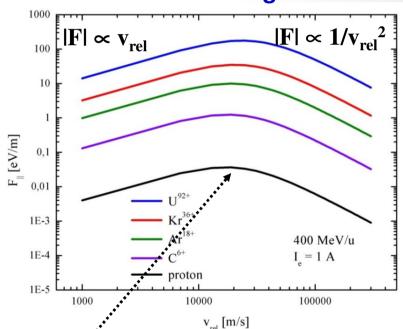


cooling force F

for small relative velocity: ∝ v_{rel}

for large relative velocity: $\propto v_{rel}^{-2}$

increases with charge: ∝ Q²



maximum of cooling force at effective electron temperature

Electron Cooling Time

first estimate:
$$\tau = \frac{3}{8\sqrt{2\pi}n_{e}Q^{2}r_{e}r_{i}cL_{C}}(\frac{k_{B}T_{e}}{m_{e}c^{2}} + \frac{k_{B}T_{i}}{m_{i}c^{2}})^{3/2}$$
 (Budker 1967)

for large relative velocities

large relative velocities cooling time
$$\tau_z \propto \frac{A}{Q^2} \frac{1}{n_e \eta} \beta^3 \gamma^5 \theta_z^3$$

$$\theta_{x,y} = \frac{v_{x,y}}{\gamma \beta c}$$

$$\theta_{\parallel} = \frac{v_{\parallel}}{\gamma \beta c}$$

cooling rate (τ^{-1}) :

- slow for hot hadron beams $\propto \theta^{-3}$
- decreases with energy $\propto \gamma^2 (\beta \cdot \gamma \cdot \theta)$ is conserved)
- linear dependence on electron beam intensity n_e and cooler length η=L_{ec}/C
- favorable for highly charged ions Q²/A
- independent of hadron beam intensity

for small relative velocities

cooling rate is constant and maximum at small relative velocity

$$F \propto v_{rel} \Rightarrow \tau = \Delta t = p_{rel}/F = constant$$

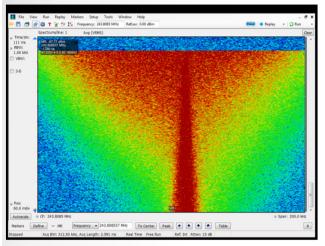
Longitudinal Cooling

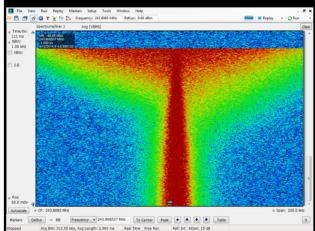
Xe⁵⁴⁺ 350 MeV/u

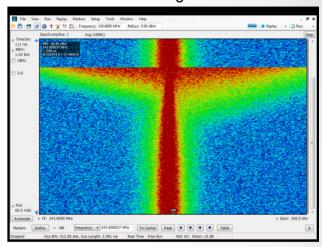
 $I_e = 100 \text{ mA}$

 $I_e = 250 \text{ mA}$

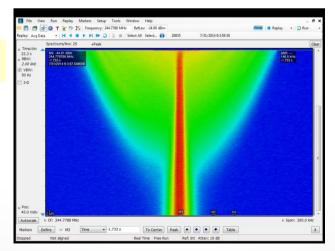
 $I_e = 500 \text{ mA}$







protons 400 MeV (Q=1)



measurement time 20 s

measurement time 650 s

 I_e = 250 mA

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Electron Beam Properties

electron beam temperature

is determined by the thermal cathode temperature k_BT_{cat}

transverse temperature $k_B T_{\perp} = k_B T_{cat}$,

can be reduced by transverse magnetic expansion with ($\propto B_c/B_{gun}$)

longitudinal temperature $k_B T_{\parallel} = (k_B T_{cat})^2/4E_0 << k_B T_{\perp}$

lower limit :
$$k_B T_{\parallel} \geq 2e \frac{n_e^{1/3}}{4\pi\epsilon_0}$$

typical values:

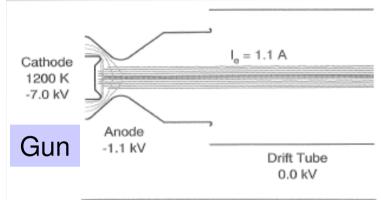
transverse $k_BT_{\perp} \approx 100 \text{ meV} (1100 \text{ K})$

with magnetic expansion $k_BT_1 \approx 1 \text{ meV}$

 $longitudinal \qquad \qquad k_B T_{\parallel} \approx 0.1 \text{ - 1 meV}$

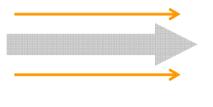
Electron Beam Properties

constant electron beam radius

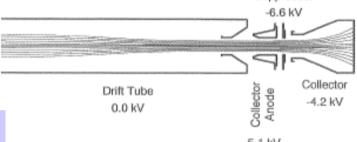


electron beam confined by longitudinal magnetic field (from gun to collector)





Cooling Section



Suppressor

transversely expanded electron beam

electron current (space charge limited)

$$I_e = PU_{an}^{3/2}$$

O Contract C

radial variation of electron energy due to space charge

$$E(r) = eU_{cat} - n_e \pi r_0^2 r_e m_e c^2 [1 + 2 \ln (r_{tube}/r_0)] + n_e \pi r_e m_e c^2 r^2$$

Electron Motion in Longitudinal Magnetic Field

single particle cyclotron motion

cyclotron frequency
$$\omega_c = \frac{eB}{\gamma me}$$

cyclotron frequency
$$\omega_c = \frac{e_B}{\gamma me}$$
 cyclotron radius $r_c = \frac{v_\perp}{\omega_c} = \frac{(kBT_\perp m_e)^{1/2} \, \gamma}{eB}$

electrons follow the magnetic field line adiabatically

⇒ transverse magnetic expansion results in a reduction of the transverse temperature

$$\frac{mv_{\perp}^{2}}{B} = const.$$

another important consequence:

for interaction times which are long compared to the cyclotron period the ions do not sense the transverse electron temperature

$$\Rightarrow$$
 magnetized cooling ($T_{eff} \approx T_{\parallel} << T_{\perp}$)

 $r_c \approx 10 \ \mu m$

Optimized Electron Cooling

minimize relative velocity between ions and electrons

electron beam space charge:

transverse electric field + longitudinal B-field ⇒ azimuthal drift

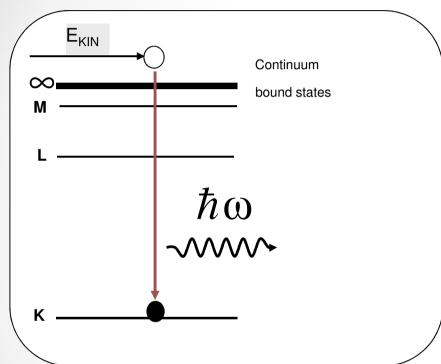
$$v_{azi} = r\omega_{azi} = r\frac{2\pi r_e n_e c^2}{\gamma\omega_c}$$

⇒ • electron and ion beam should be well centered

Favorable for optimum cooling (small transverse relative velocity):

- parallel adjustment of ion and electron beam
- high parallelism of magnetic field lines B_⊥/B_{||} in cooling section
- large beta function (small divergence) in cooling section

Atomic Physics Limitation of Electron Cooling



Radiative Electron Capture (REC)

$$A^{Q+} + e^{-} A^{(Q-1)+} + hv$$

emission of a photon

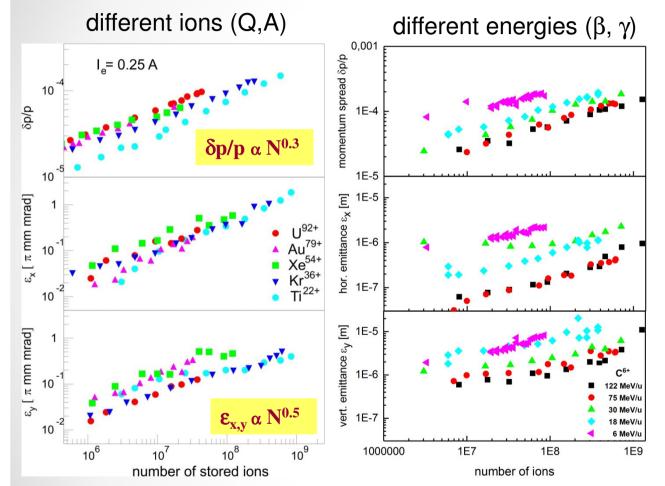
change of the ion charge results in particle loss ⇒ different orbit

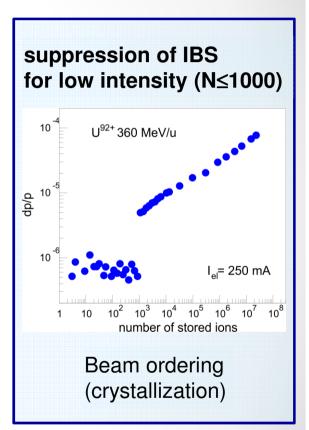
loss rate
$$\tau^{-1} = \gamma^2 \, \alpha_{\rm REC} \, n_{\rm e} \eta$$

$$\alpha_{\rm REC} = \frac{1.92 \times 10^{-13} \, Q^2}{\sqrt{k_{\rm B}T}} \, \left(ln \, \frac{5.66 \, Q}{\sqrt{k_{\rm B}T}} + 0.196 (\frac{k_{\rm B}T}{Q^2})^{1/3} \right) [cm^3 \, s^{-1}]$$

losses by recombination (REC)

Electron Cooled Beams in Equilibrium with Intrabeam Scattering (IBS)





heating rate dominated by Intrabeam Scattering

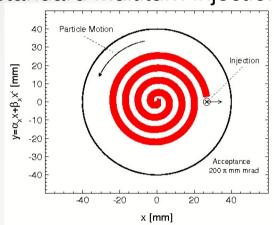
$$\tau_{IBS}^{-1} = \frac{Q^4 e^4}{\left(Am_i^2\right)^2} \cdot \frac{N}{C\varepsilon_h \varepsilon_v \delta p / p} \cdot \frac{1}{\left(\gamma^4 \beta^3 c^3\right)} \cdot 4\pi L_C^{IBS}$$

IBS: total phase space volume increases with ion beam intensity and ion charge

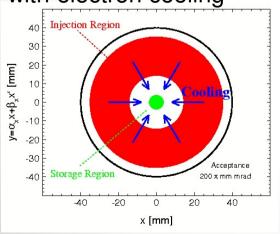
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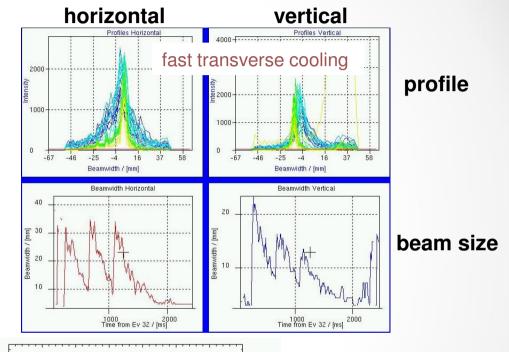
Accumulation of Heavy Ions by Electron Cooling

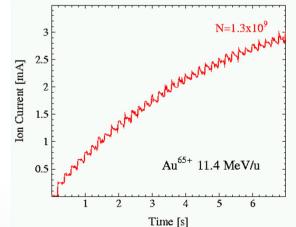
standard multiturn injection



fast accumulation by repeated multiturn injection with electron cooling







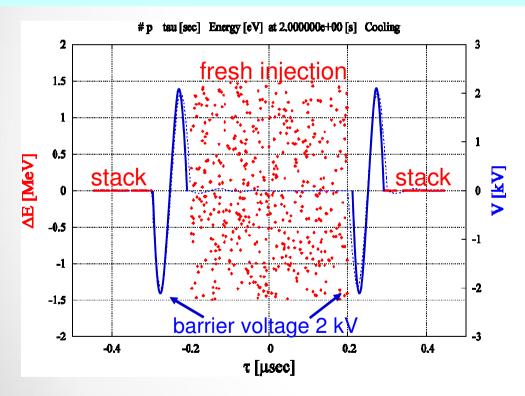
intensity increase in 5 s by a factor of \approx 10

limitations: space charge tune shift, recombination (REC)

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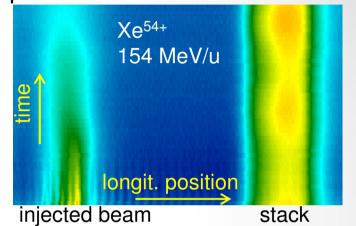
Accumulation of Secondary Particles

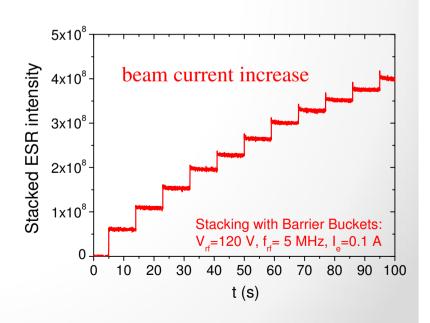
basic idea: confine stored beam to a fraction of the circumference, inject into gap and apply cooling to merge the two beam components ⇒ fast increase of intensity (for secondary beams)



simulation of longitudinal stacking with barrier buckets and electron cooling

experimental verification at ESR

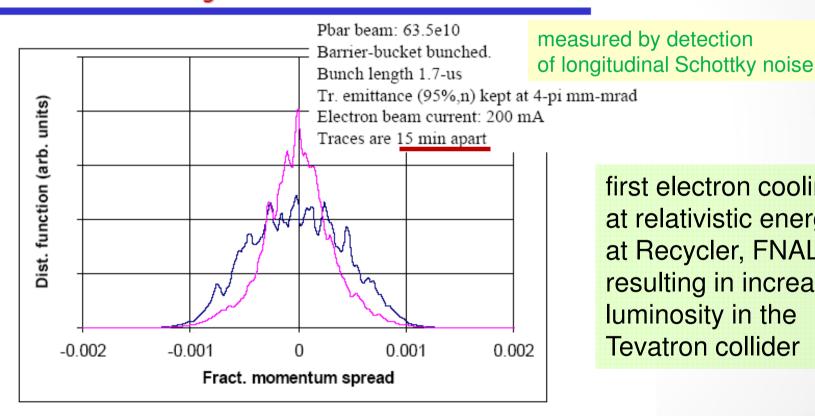




High Energy Electron Cooling

electron cooling of 8 GeV antiprotons longitudinal cooling with 0.2 A, 4.4 MeV electron beam

First e-cooling demonstration - 07/15/05

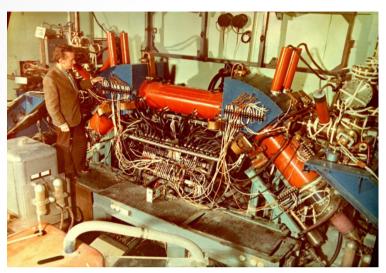


first electron cooling at relativistic energy at Recycler, FNAL resulting in increased luminosity in the Tevatron collider

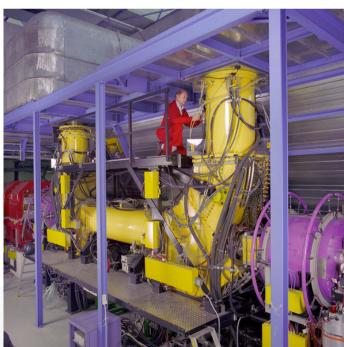
cooling time of some ten minutes has to be compared with the accumulation time of many hours

Electron Cooling Systems

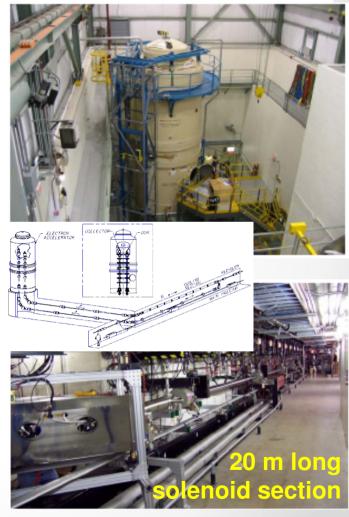
First Electron Cooling System NAP-M/BINP 1974



Medium Energy: 300 keV ESR/GSI 1990



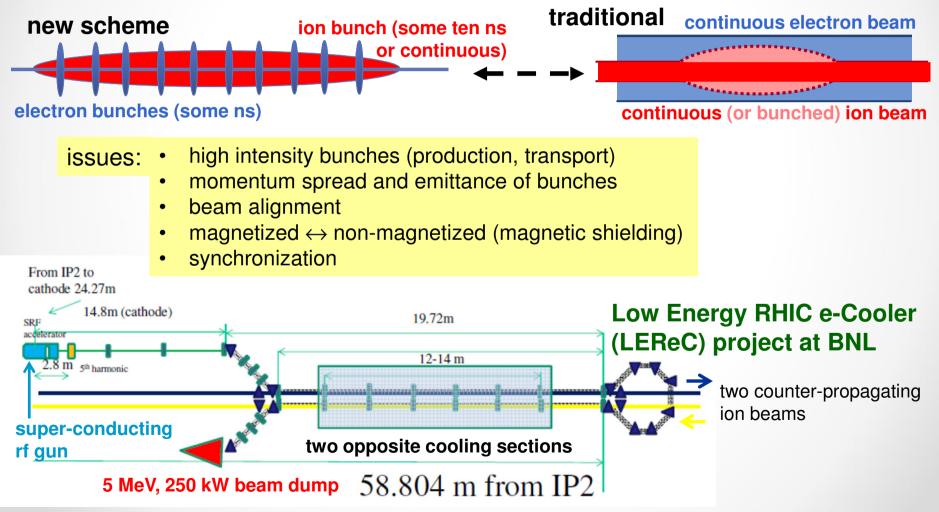
High Energy: 4.3 MeV Recycler/FNAL 2005



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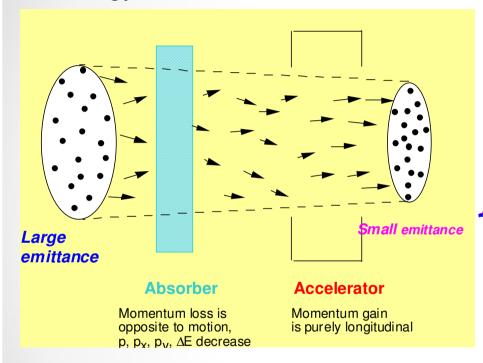
Bunched Beam Electron Cooling

Electron cooling with electrostatic acceleration is limited in energy (5-10 MeV). A bunched electron beam offers the extension of the electron cooling method to higher energy (linear rf accelerator).

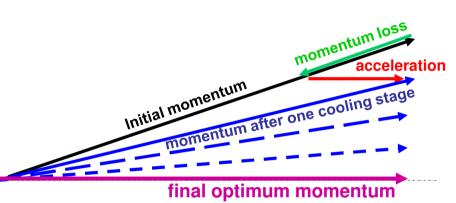


2. Ionization Cooling

energy loss in solid matter



proposed for muon cooling



not useful for heavy particles due to strong interaction with matter

transverse cooling

$$\frac{d\epsilon_N}{ds} = -\frac{1}{\beta^2 E} \frac{dE}{ds} \epsilon_N + \frac{\beta \gamma \beta_\perp}{2} \frac{\langle \theta_{rms}^2 \rangle}{ds}$$
$$= -\frac{1}{\beta^2 E} \frac{dE}{ds} \epsilon_N + \frac{\beta_\perp E_s^2}{2\beta^3 m_\mu c^2 L_R E}$$

 \Rightarrow small β_{\perp} at absorber in order to minimize multiple scattering

large L_R , $(dE/ds) \Rightarrow light absorbers <math>(H_2)$

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Ionization Cooling

increased longitudinal cooling by longitudinal-transverse emittance exchange

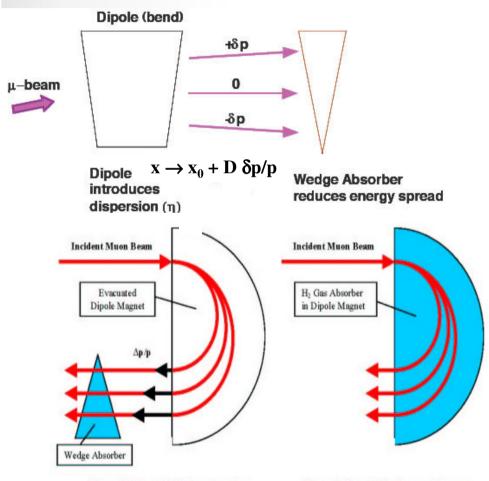


Figure 1. Use of a Wedge Absorber for Emittance Exchange

Figure 2. Use of Continuous Gaseous Absorber for Emittance Exchange

$$\frac{d\sigma_E^2}{ds} = -2\frac{\partial (dE/ds)}{\partial E}\sigma_E^2 + \frac{d\langle \Delta E_{rms}^2\rangle}{ds}$$
 cooling term heating term

cooling, if
$$\frac{\partial (dE/ds)}{\partial E} > 0$$

emittance exchange

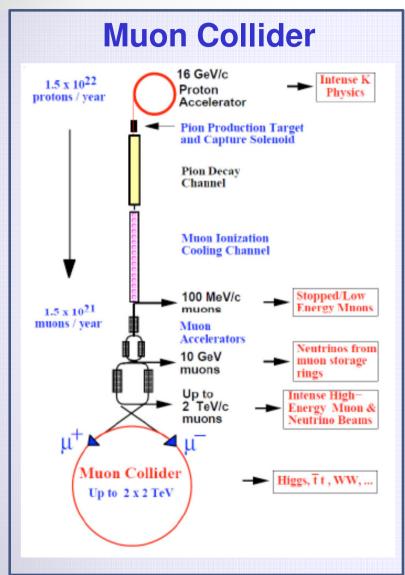
increased longitudinal cooling

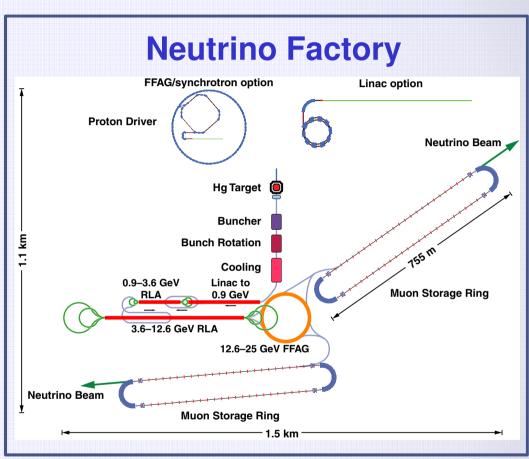
$$\frac{\partial \frac{dE}{ds}}{\partial E} \Rightarrow \frac{\partial \frac{dE}{ds}}{\partial E}|_{0} + \frac{dE}{ds} \frac{D\rho'}{\beta c p \rho_{0}}$$

reduced transverse cooling

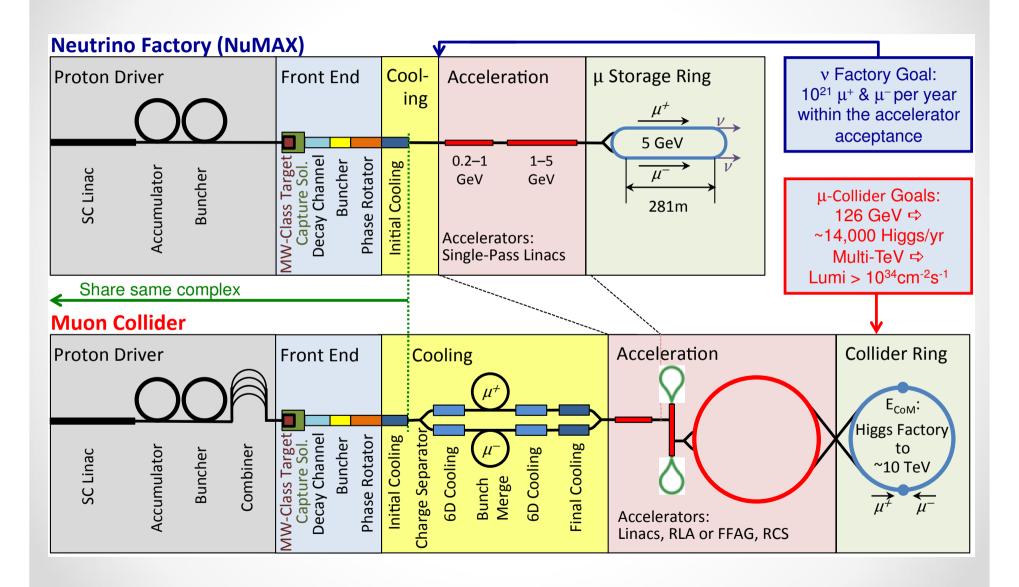
$$\frac{d\epsilon_N}{ds} = -\frac{1}{\beta^2 E} \frac{dE}{ds} (1 - \frac{D\rho'}{\rho_0}) \epsilon_N$$

Scenarios with Ionization Cooling



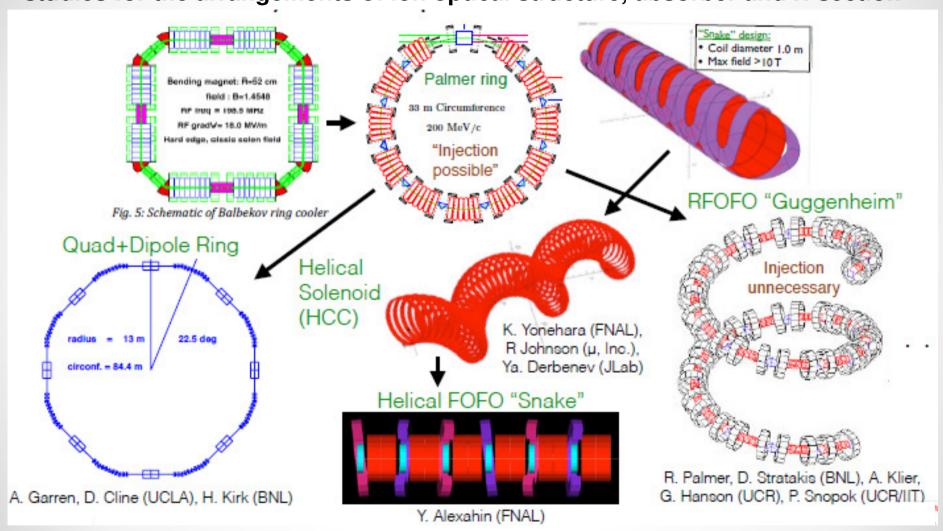


Scenarios with Ionization Cooling

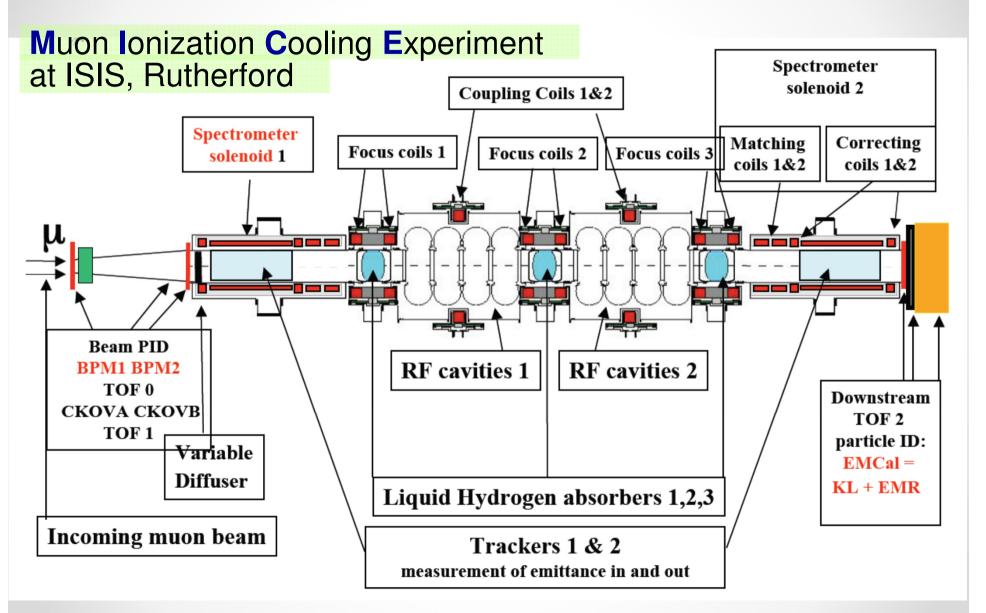


The Muon Cooling Section

studies for the arrangements of ion optical structure, absorber and rf section

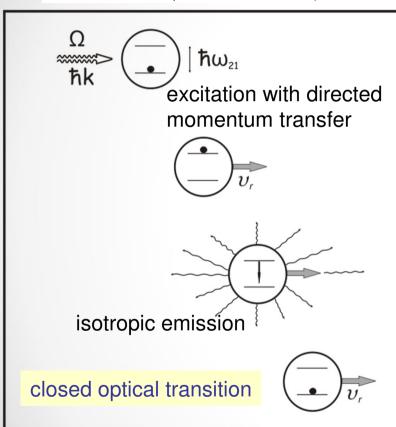


MICE

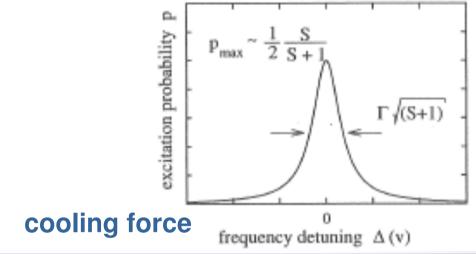


3. Laser Cooling

$$\Omega = \gamma \omega_{21} (1 - \beta \cos \theta)$$



the directed excitation and isotropic emission result in a transfer of velocity v_r



$$\overrightarrow{F}(\overrightarrow{v}, \overrightarrow{k}) = \frac{\hbar \overrightarrow{k}}{2} S\Gamma \frac{(\Gamma/2)^2}{(\omega - \omega_{21} - \overrightarrow{v} \overrightarrow{k}) + (\Gamma/2)^2 (1+S)}$$

Lorentzian with width $\Gamma/k \sim 10 \text{ m/s}$

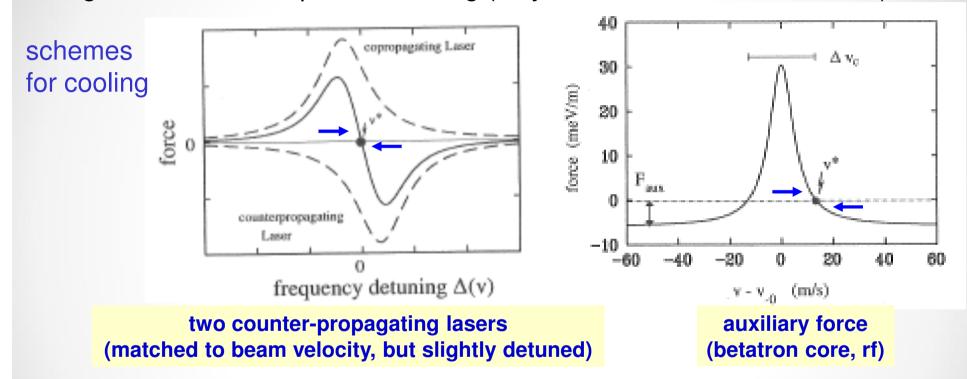
minimum temperature
$$\,T_D=\frac{\hbar\Gamma}{2k_B}\,$$
 (Doppler limit) typical 10⁻⁵ – 10⁻⁴ K

typical cooling time ~ 10 μs

drawback: only longitudinal cooling

Laser Cooling

a single laser does not provide cooling (only acceleration or deceleration)

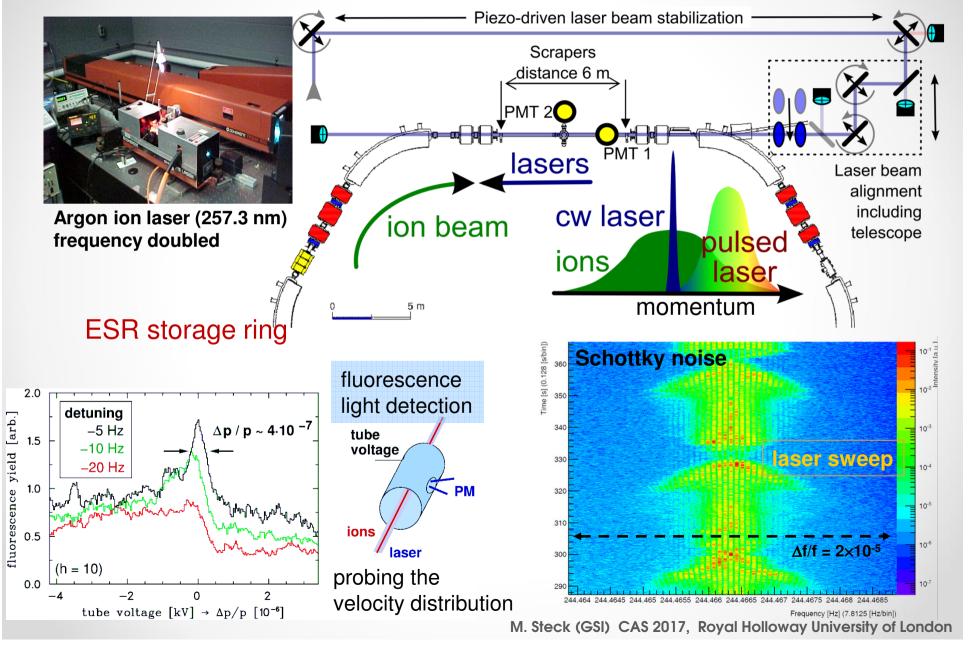


capture range of laser is limited ⇒ frequency sweep (snowplow) or pulsed laser with large spectral width

ions studied so far: ⁷Li¹⁺, ⁹Be¹⁺, ²⁴Mg¹⁺, ¹²C³⁺

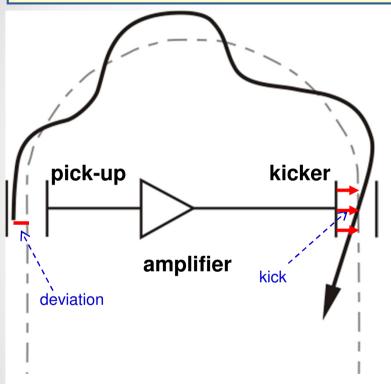
in future: Li-like heavy ions at relativistic energies, cooling rate increases with γ large relativistic energy \Rightarrow large excitation energy in PRF

Laser Cooling of C³⁺



4. Stochastic Cooling

First cooling method which was successfully used for beam preparation



S. van der Meer, D. Möhl, L. Thorndahl et al. (1925 – 2011) (1936-2012)

Conditions:

Betatron motion phase advance (pick-up to kicker): $(n + \frac{1}{2}) \pi$

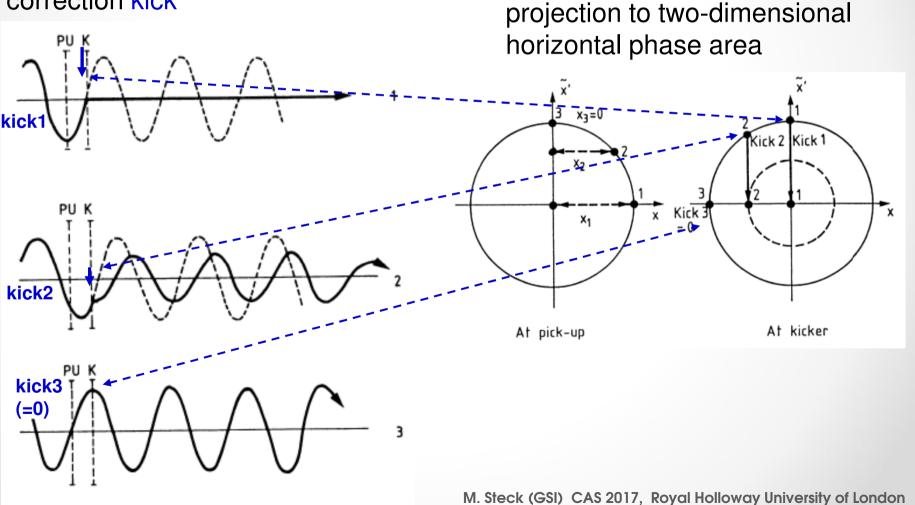
Signal travel time = time of flight of particle (between pick-up and kicker)

Sampling of sub-ensemble of total beam

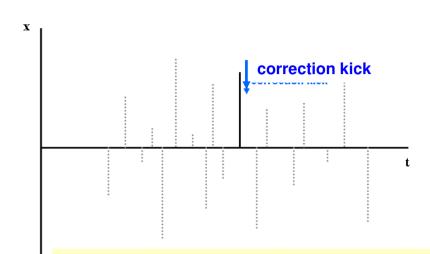
Principle of transverse cooling: measurement of deviation from ideal orbit is used for correction kick (feedback)

Stochastic Cooling

single particle betatron motion along storage ring without (dashed) and with (full) correction kick



Stochastic Cooling

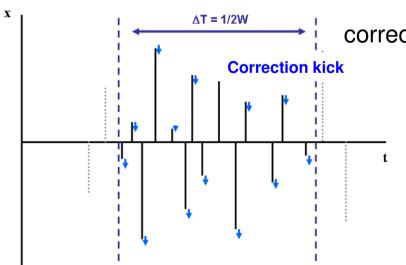


in time domain

correction kick (unlimited resolution)

$$\Delta x = g \times x$$

Nyquist theorem: a system with a band-width $\Delta f = W$ in frequency domain can resolve a minimum time duration $\Delta T = (2W)^{-1}$



correction kick
$$\Delta x = \frac{g}{N_s} \times \sum_{i=1..N_s} x_i$$
 , $N_s = N \frac{\Delta T}{T_0} = \frac{N}{2WT_0}$

For exponential damping $(x(t) = x(t_0) \times e^{-(t-t_0)/\tau})$:

$$\tau^{-1} = T_0^{-1} \times \frac{\Delta x}{x} = \frac{g2W}{N}, \ if \sum_{i=1..N_s} x_i = x$$

cooling rate

$$\tau^{-1} \le \frac{2W}{N} \quad if \ g \le 1$$

Stochastic Cooling

some refinements of cooling rate formula

noise: thermal or electronic noise adds to the beam signal

mixing: change of relative longitudinal position of particles due to momentum spread

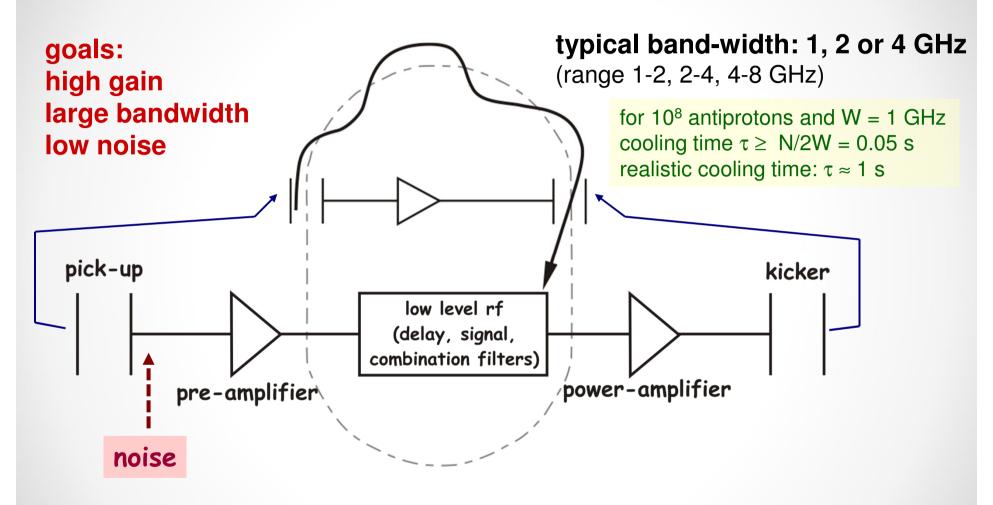
cooling rate
$$\lambda=\tau^{-1}=\frac{2W}{N}(\underline{2g}-\underline{g^2(M+U)})$$
 M mixing factor U noise to signal ratio heating

$$\frac{d\lambda}{dg} = 0 \Rightarrow g = \frac{1}{M+U}$$

further refinement (wanted \leftrightarrow unwanted mixing):

with wanted mixing M (kicker to pick-up) $\lambda = \tau^{-1} = \frac{2W}{N}(2g(1-\tilde{M}^2) - g^2(M+U))$ and unwanted mixing \tilde{M} (pick-up to kicker)

Stochastic Cooling Circuit



Transfer Function:

$$Z_{pick-up} \cdot G_{pick-up}(E) \cdot H(t_{delay}) \cdot F(E) \cdot G \cdot G_{kicker}(E) \cdot Z_{kicker}$$

Longitudinal Stochastic Cooling

1) Palmer cooling

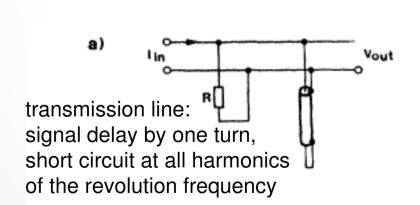
pick-up in dispersive section detects horizontal position

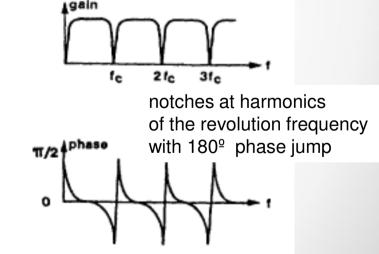
⇒ acceleration/deceleration kick corrects momentum deviation

2) Notch filter cooling

filter creates notches at the harmonics of the nominal revolution frequency

⇒ particles are forced to circulate at the nominal frequency



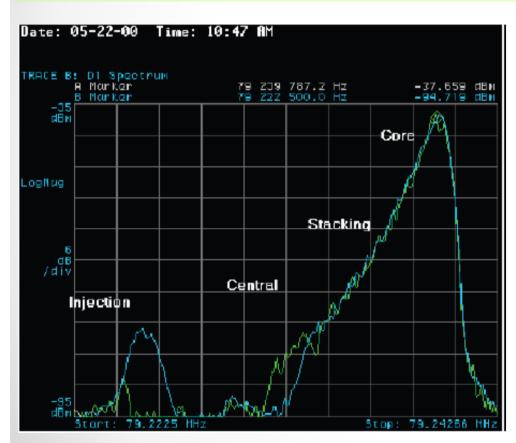


3) ToF cooling

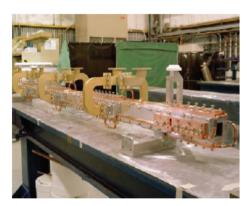
simplified scheme without notches allows efficient pre-cooling

Antiproton Accumulation by Stochastic Cooling

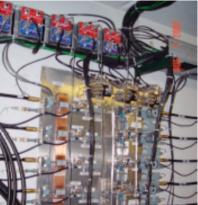
accumulation of 8 GeV antiprotons at accumulator ring, FNAL, shut down 09/2011 a similar facility AC/AA at CERN was operated until 11/1996



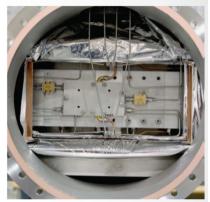
momentum distribution of accumulated antiproton beam



kicker array



microwave electronics



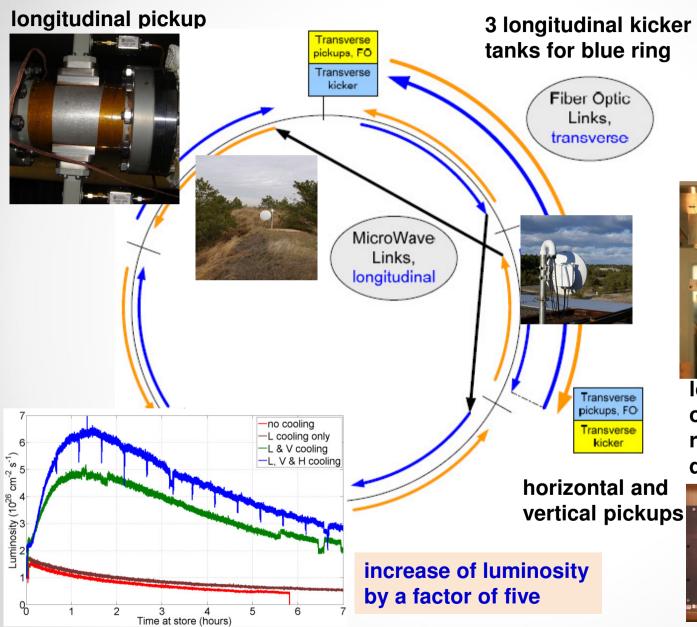
cryogenic microwave amplifier



power amplifiers (TWTs)

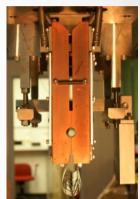
M. Steck (GSI) CAS 2017, Royal Holloway University of London

RHIC – 3D stochastic cooling for heavy ions

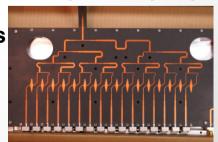








longitudinal kicker open for injection and ramping (left), closed during cooling (right)



Stochastic Cooling of Rare Isotopes at GSI

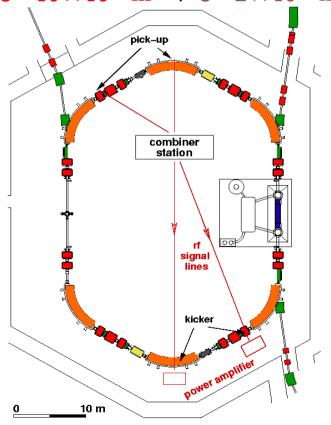
fast pre-cooling of hot fragment beams

energy 400 (-550) MeV/u

bandwidth 0.8 GHz (range 0.9-1.7 GHz)

$$δp/p = ±0.35 % → δp/p = ±0.01 %$$

 $ε = 10 × 10-6 m → ε = 2 × 10-6 m$





electrodes installed inside magnets



combination of signals from electrodes



power amplifiers for generation of correction kicks

M. Steck (GSI) CAS 2017, Royal Holloway University of London

Comparison of Cooling Methods

Stochastic Cooling Electron Cooling

Useful for: low intensity beams low energy

all intensities

hot (secondary) beams warm beams (pre

high charge

full 3D control

warm beams (pre-cooled)

high charge

bunched beams

Limitations: high intensity beams

/problems beam quality limited

bunched beams

space charge effects

recombination losses

high energy

laser cooling (of incompletely ionized ions) and ionization cooling (of muons) are quite particular and not general cooling methods

References 1 (general)

- A. Chao, M. Tigner, Handbook of Accelerator Physics and Engineering, Chapter 2.8, World Scientific, Singapore, 1999
- M. Minty, F. Zimmermann, Measurement and Control of Charged Particle Beams, Chapter 11, Springer Verlag, Berlin, 2003
- D. Möhl, Principle and Technology of Beam Cooling, CERN/PS 86-31,1986
- D. Möhl, Beam Cooling, CAS 2005, CERN 2005-04, pp.324-339
- H. Danared, Beam Cooling CAS 2005, CERN 2005-06, pp. 343-362
- Y. Zhang, W. Chou, ICFA Beam Dynamics Newsletter No. 64 and 65, December 2014, http://icfa-usa.jlab.org/archive/newsletter.shtml

References 2 (specialized)

Electron Cooling:

- H. Poth, Electron Cooling, CAS 85, CERN 87-03, pp. 534-569, 1987
- H. Poth, Electron Cooling: Theory, Experiment, Application, Phys. Rep. Vol. 196 Issues 3-4, pp. 135-297, 1990
- I. Meshkov, Electron Cooling: Status and Perspectives, Physics of Particles and Nuclei, Vol. 25, Issue 6, pp. 631-661, 1994

Stochastic Cooling:

- D. Möhl, Stochastic Cooling for Beginners, CAS 1983, CERN 84-15, pp. 97-162
- D. Möhl, Stochastic Cooling, CAS 85, CERN 87-03, pp. 453-533, 1987
- D. Möhl, Stochastic Cooling of Particle Beams, Springer Lecture Notes in Physics 866 (2013)
- S. van der Meer, Rev. Mod. Phys. Vol. 57, No. 3 Part 1, 1985

Laser Cooling:

E. Bonderup, Laser Cooling, CAS 1993, CERN 95-06, pp. 731-748

Ionization Cooling:

D. Neuffer, Introduction to Muon Cooling, Nucl. Instr. Meth. A 532 (2004) 26-31

Proceedings of Biannual Workshops on Beam Cooling: e.g. COOL 2017, Bonn, Germany